Magnetic field structure in single late-type giants: The fast rotator FI Cnc

Svetla Tsvetkova*¹, Pascal Petit², Renada Konstantinova-Antova¹, Ana Borisova¹, Michel Auriere², and Corinne Charbonnel³

¹Institute of Astronomy and NAO, BAS (IA and NAO) – 72 Tsarigradsko shosse, Sofia, Bulgaria
 ²Institut de recherche en astrophysique et planétologie (IRAP) – Université Paul Sabatier - Toulouse 3,
 Observatoire Midi-Pyrénées, Centre National de la Recherche Scientifique : UMR5277 – France
 ³Department of Astronomy, University of Geneva – Chemin des Maillettes 51, 1290 Versoix,
 Switzerland, Switzerland

Abstract

We present a study of the 2.4 Msun giant FI Cnc, which is at the base of the red giant branch. We use spectropolarimetric data obtained in two sets (2013 and 2015) with Narval @ TBL, Pic du Midi Observatory, France. Employing ZDI technique, we are able to reconstruct the first map of the global surface magnetic field structure of the star. In addition, we measure the line activity indicators and Bl (average longitudinal magnetic field via LSD method) and compare their behavior with time. FI Cnc shows clear evidences that its magnetism is due to dynamo.

^{*}Speaker